

Hartree-Fock

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1 Generalities

Given a set of orthonormal spin-orbitals $\{\psi_\mu(\mathbf{x})\}_{\mu=1}^N$ (where $\mathbf{x} = (\mathbf{r}, \sigma)$ represents the spatial-spin coordinate of an electron)

$$\int \psi_\mu^*(\mathbf{x}) \psi_\nu(\mathbf{x}) d^4\mathbf{x} = \delta_{\mu\nu}$$

a **Slater determinant** can be defined as

$$\Psi(\mathbf{x}_1, \dots, \mathbf{x}_N) = \begin{vmatrix} \psi_1(\mathbf{x}_1) & \psi_1(\mathbf{x}_2) & \dots & \psi_1(\mathbf{x}_N) \\ \psi_2(\mathbf{x}_1) & \psi_2(\mathbf{x}_2) & & \psi_2(\mathbf{x}_N) \\ \dots & \dots & & \dots \\ \psi_N(\mathbf{x}_1) & \psi_N(\mathbf{x}_2) & \dots & \psi_N(\mathbf{x}_N) \end{vmatrix}$$

Alternatively, in abstract notation¹, if $\{|\mu\rangle\}_{\mu=1}^N$ denotes the set of spin-orbitals we write the orthonormalization condition as

$$\langle \mu | \nu \rangle = \delta_{\mu\nu}$$

¹In the following we consider a set of states which we label with a set of integers $|1\rangle, |2\rangle, \dots, |N\rangle$ rather than using the more cumbersome notation $|\psi_1\rangle, |\psi_2\rangle, \dots, |\psi_N\rangle$.

and the Slater determinant as

$$\begin{aligned} |\Psi\rangle &= \frac{1}{\sqrt{N!}} \sum_{P \in S_N} \epsilon_P P \{|1\rangle |2\rangle \dots |N\rangle\} \\ &= \sqrt{N!} A_N \{|1\rangle |2\rangle \dots |N\rangle\} \end{aligned}$$

Here P is a **permutation operator** ($P^\dagger = P^{-1}$) and runs over the group S_N of permutations of the first N integers²

$$\begin{aligned} P &= \begin{pmatrix} 1 & 2 & \dots & N \\ p_1 & p_2 & \dots & p_N \end{pmatrix} \Rightarrow \\ P \{|1\rangle |2\rangle \dots |N\rangle\} &= \{|p_1\rangle |p_2\rangle \dots |p_N\rangle\} \end{aligned}$$

The **antisymmetrizer** A_N introduced above has been defined according to

$$A_N = \frac{1}{N!} \sum_{P \in S_N} \epsilon_P P$$

(i.e. as a linear combination of the P operators with coefficients given by the *parity* ϵ_P) and satisfies $A_N^\dagger = A_N = A_N^2$. Hence, A_N defined in this way is a *projector*.

The wavefunction is correctly normalized. Indeed, using the projection properties of A_N we obtain

$$\begin{aligned} \langle \Psi | \Psi \rangle &= N! \langle A_N(12\dots N) | A_N(12\dots N) \rangle = \\ &= N! \langle 12\dots N | A_N^2 | 12\dots N \rangle \\ &= N! \langle 12\dots N | A_N | 12\dots N \rangle \end{aligned}$$

and

$$N! \langle 12\dots N | A_N | 12\dots N \rangle = N! \frac{1}{N!} \sum_{P \in S_N} \epsilon_P \langle 12\dots N | p_1 p_2 \dots p_N \rangle \equiv 1$$

where in the last equality $\langle 12\dots N | p_1 p_2 \dots p_N \rangle = \delta_{1p_1} \delta_{2p_2} \dots \delta_{Np_N}$ has been used. More generally, for $|\Psi\rangle$ and $|\Psi'\rangle$ defined as

$$|\Psi\rangle = \sqrt{N!} A_N |\Phi\rangle \quad \text{and} \quad |\Psi'\rangle = \sqrt{N!} A_N |\Phi'\rangle$$

²The notation for P means that 1 must be replaced by p_1 , 2 by p_2 and so on. Thus p_1, p_2, \dots, p_N is the same set of integers on the first line, but possibly with a different order. The whole set of P makes up a group, called the **symmetric group** - of order N - and usually denoted as S_N .

we have

$$\langle \Psi | \Psi' \rangle = N! \langle \Phi | A_N | \Phi' \rangle = \sum_{P \in S_N} \epsilon_P \langle \Phi | P | \Phi' \rangle$$

This is actually a special case of a general result: for any operator O which is *symmetric* under exchange of the electron labels it holds

$$[P, O] = [A_N, O] = 0$$

hence

$$\langle \Psi | O | \Psi' \rangle = \sum_{P \in S_N} \epsilon_P \langle \Phi | O P | \Phi' \rangle$$

These formula replace the $N!$ sum appearing in the scalar product between two antisymmetrized vectors with a simpler $N!$ sum.

2 Expectation values

Let us now write down the expectation value of the Hamiltonian operator on a Slater determinant

$$\langle \Psi | H | \Psi \rangle = E_{HF} \quad H = \sum_{i=1}^N h_i + \frac{1}{2} \sum_{j, i \neq j}^N g_{ij} = H^{(1)} + H^{(2)}$$

Here h_i 's are monoelectronic operators, $h_i = k_i + v_i$, and g_{ij} is the interelectronic repulsion potential. Correspondingly, $H^{(1)}$ and $H^{(2)}$ are the one- and two- electron terms of the Hamiltonian, respectively. Both $H^{(1)}$ and $H^{(2)}$ are symmetric operators, hence

$$\begin{aligned} \langle \Psi | H^{(i)} | \Psi \rangle &= N! \langle A_N(12..N) | H^{(i)} | A_N(12..N) \rangle = \\ &= N! \langle 12..N | A_N H^{(i)} A_N | 12..N \rangle = \\ &= N! \langle 12..N | H^{(i)} A_N^2 | 12..N \rangle = \\ &= \sum_{P \in S_N} \epsilon_P \langle 12..N | H^{(i)} | p_1 p_2 .. p_N \rangle \end{aligned}$$

Let us first consider the monoelectronic term and focus attention on the first electron

$$\begin{aligned} \sum_{P \in S_N} \epsilon_P \langle 12..N | h_1 | p_1 p_2 .. p_N \rangle &= \\ = \sum_{P \in S_N} \epsilon_P \langle 1 | h_1 | p_1 \rangle \langle 2 | p_2 \rangle .. \langle N | p_N \rangle &= \\ = \sum_{P \in S_N} \epsilon_P \langle 1 | h_1 | p_1 \rangle \delta_{2p_2} .. \delta_{Np_N} &= \langle 1 | h | 1 \rangle \end{aligned}$$

In the last line the subscript 1 on h has been removed since it is dummy. By the same token we have

$$\sum_{P \in S_N} \epsilon_P \langle 12..N | h_\mu | p_1 p_2 .. p_N \rangle = \langle \mu | h | \mu \rangle$$

Thus, we can write the one-electron Hamiltonian as

$$\langle \Psi | H^{(1)} | \Psi \rangle = \sum_{\mu=1}^N \langle \mu | h | \mu \rangle$$

Analogously

$$\begin{aligned} & \sum_{P \in S_N} \epsilon_P \langle 12..N | g_{12} | p_1 p_2 .. p_N \rangle = \\ &= \sum_{P \in S_N} \epsilon_P \langle 12 | g_{12} | p_1 p_2 \rangle \delta_{3p_3} \delta_{4p_4} .. \delta_{Np_N} = \\ &= \langle 12 | g | 12 \rangle - \langle 12 | g | 21 \rangle \end{aligned}$$

i.e. in general

$$\langle \Psi | H^{(2)} | \Psi \rangle = \frac{1}{2} \sum_{\mu, \nu (\neq \mu)=1}^N (\langle \mu \nu | g | \mu \nu \rangle - \langle \mu \nu | g | \nu \mu \rangle)$$

This can be rewritten in a simpler form by introducing an **exchange operator** π ³

$$\begin{aligned} \pi | \mu \nu \rangle &= | \nu \mu \rangle \\ \langle \Psi | H^{(2)} | \Psi \rangle &= \frac{1}{2} \sum_{\mu, \nu (\neq \mu)=1}^N \langle \mu \nu | g(1 - \pi) | \mu \nu \rangle \end{aligned}$$

In conclusion, the expectation value of H (the **Hartree-Fock energy**) reads as

$$\langle \Psi | H | \Psi \rangle = \sum_{\mu=1}^N \langle \mu | h | \mu \rangle + \frac{1}{2} \sum_{\mu, \nu (\neq \mu)=1}^N \langle \mu \nu | g(1 - \pi) | \mu \nu \rangle$$

Note that in the last term on the r.h.s. we can remove the restriction $\nu \neq \mu$ since the term $\nu = \mu$ does not contribute to the sum.

³It can be readily shown that $\pi = \pi^\dagger = \pi^{-1} = \pi^2$ and $\pi g = g\pi$.

3 Variation of the orbitals

Let us now apply the variational principle. In doing this, we take care of the orthogonalization condition

$$\langle \mu | \nu \rangle = 1$$

which implies

$$\langle \delta\mu | \nu \rangle + \langle \mu | \delta\nu \rangle = 0$$

or, equivalently,

$$\begin{aligned} \langle \delta\mu | \nu \rangle &= 0 & \mu &\neq \nu \\ \Re \langle \delta\mu | \mu \rangle &= 0 & \mu &= \nu \end{aligned}$$

and focus on the (special class of) variations that satisfy⁴

$$\langle \delta\mu | \nu \rangle = 0 \quad \text{for any } \nu = 1, 2, \dots, N$$

From a geometrical point of view they satisfy $|\delta\mu\rangle \in \{|\nu\rangle\}_{\nu=1..N}^\perp$, i.e. they are such that the $|\delta\mu\rangle$ is (arbitrarily) confined in the orthogonal complement of the “occupied” orbital space⁵. Applying the variation of the orbital $|\mu\rangle$ we obtain

$$\begin{aligned} 0 = \langle \delta\Psi | H\Psi \rangle &= \langle \delta\mu | h | \mu \rangle + \\ &+ \frac{1}{2} \sum_{\nu(\neq\mu)=1}^N \langle \delta\mu\nu | g(1-\pi) | \mu\nu \rangle + \\ &+ \frac{1}{2} \sum_{\nu(\neq\mu)=1}^N \langle \nu\delta\mu | g(1-\pi) | \nu\mu \rangle = \\ &= \langle \delta\mu | h | \mu \rangle + \sum_{\nu(\neq\mu)=1}^N \langle \delta\mu\nu | g(1-\pi) | \mu\nu \rangle = \\ &\equiv \langle \delta\mu | h | \mu \rangle + \sum_{\nu=1}^N \langle \delta\mu\nu | g(1-\pi) | \mu\nu \rangle \end{aligned}$$

⁴Variations of this kind are complex-linear, meaning they form a *complex* linear space, something that is not possible under the more general condition $\Re \langle \delta\mu | \mu \rangle = 0$. The remaining variations add nothing, and complicate the analysis because they would require application of the stationary condition in “full form”, in contrast to the “Dirac-Frenkel” form $\langle \delta\Psi | H | \Psi \rangle = 0$.

⁵The symbol \perp denotes the **orthogonal complement**. Remember that the orthogonal complement of a set of vector - *however chosen* - is the *linear space* spanned by the vectors that are orthogonal to each vector of the given set. That is, $\{|1\rangle, |2\rangle, \dots, |m\rangle\}^\perp = \{|\psi\rangle \text{ such that } \langle \mu | \psi \rangle = 0 \text{ for } \mu = 1, m\}$.

where $\pi |\mu\mu\rangle = |\mu\mu\rangle$ has been used in the last line. Thus,

$$\langle \delta\Psi | H\Psi \rangle = \langle \delta\mu | h + \sum_{\nu=1}^N \langle \nu | g(1 - \pi) | \mu \rangle | \nu \rangle = \langle \delta\mu | f | \mu \rangle = 0 \quad (1)$$

where f is a mono-electronic, *self-adjoint* operator known as **Fock operator**

$$f = h + \sum_{\nu=1}^N \langle \nu | g(1 - \pi) | \nu \rangle \quad (2)$$

Now, an equation of the form

$$\langle \delta\mu | \phi \rangle = 0$$

with $|\delta\mu\rangle$ *arbitrary* in $\{|\mu\rangle\}_{\mu=1,\dots,N}^\perp$ implies $|\phi\rangle \in \{|\mu\rangle\}_{\mu=1,\dots,N}^{\perp\perp}$, i.e., in other words,

$$f |\mu\rangle = \sum_{\nu=1}^N |\nu\rangle \epsilon_{\nu\mu}$$

where $\epsilon_{\nu\mu} = \langle \nu | f | \mu \rangle$ is a hermitean matrix since f is self-adjoint.

The Fock operator depends on the kets $|\mu\rangle$ entering the Slater determinant but, on inspecting of Eq.(2), it is clear that it is manifestly invariant under unitary transformations of the orbitals. Thus, a unitary transformation in the occupied space can be performed to put the variational equations above in “canonical” form

$$f |\mu\rangle = \epsilon_\mu |\mu\rangle \quad (3)$$

Here, the solution orbitals and eigenvalues are known as **canonical** orbitals and orbital energies.

Note 1 As noticed above, the f operator depends on $\{|\mu\rangle\}$, i.e. the HF equations are *not* simple linear equations. However, at convergence, f is a well defined operator and the $|\mu\rangle$ ’s are its eigenvectors. A closer look at f reveals that

$$f = k + v + \sum_{\nu} (\langle \nu | g | \nu \rangle - \langle \nu | g \pi | \nu \rangle)$$

or, in coordinate representation⁶,

$$\hat{f} = \hat{k} + v(\mathbf{r}) + v_H(\mathbf{r}) - \hat{v}_{\text{ex}}$$

⁶In the following the spin-orbitals are taken of the “spin-collinear” form, $\psi_\nu(\mathbf{x}) = \phi_\nu(\mathbf{r})\theta_\nu(\sigma)$, where ϕ is the spatial component and θ its spin component. Hence, ϕ_ν stands for the spatial orbital of the ν -th one-electron state employed.

where $\hat{k} = -\frac{1}{2}\nabla^2$ is the kinetic energy, $v(\mathbf{r})$ is the “external” potential (e.g. the electrostatic potential nuclear generated by the nuclei), and the remaining terms take the form

$$v_H(\mathbf{r}) = \sum_{\nu} \langle \nu | g | \nu \rangle = \int \frac{n(\mathbf{r}')}{\|\mathbf{r} - \mathbf{r}'\|} d^3\mathbf{r}' \quad \hat{v}_{\text{ex}} = \sum_{\nu} \langle \nu | g\pi | \nu \rangle$$

Here

$$n(\mathbf{r}) = \sum_{\nu} |\phi_{\nu}(\mathbf{r})|^2$$

is the **electron density** of the N electrons in the N HF orbitals, and v_H is the so-called **Hartree potential**, *i.e.* the potential felt by an electron in the field of the nuclei *and* of an N -electron distribution described by⁷ $n(\mathbf{r})$. The term v_{ex} is a *non-local exchange operator* whose matrix elementes in coordinate representation are of the form⁸

$$\langle \mathbf{x} | v_{\text{ex}} | \mathbf{x}' \rangle = \sum_{\nu} \frac{\psi_{\nu}(\mathbf{x}) \psi_{\nu}^*(\mathbf{x}')}{\|\mathbf{r} - \mathbf{r}'\|}$$

as can be seen by introducing the spectral representation of g

$$g = \int d\mathbf{x}_2 d\mathbf{x}_3 \frac{|\mathbf{x}_2 \mathbf{x}_3\rangle \langle \mathbf{x}_2 \mathbf{x}_3|}{\|\mathbf{r}_2 - \mathbf{r}_3\|}$$

and noticing that

$$\langle \mathbf{x} \nu | \mathbf{x}_2 \mathbf{x}_3 \rangle \langle \mathbf{x}_2 \mathbf{x}_3 | \nu \mathbf{x}' \rangle = \langle \mathbf{x} | \mathbf{x}_2 \rangle \phi_{\nu}^*(\mathbf{x}_3) \phi_{\nu}(\mathbf{x}_2) \langle \mathbf{x}_3 | \mathbf{x}' \rangle$$

Note 2 Let \bar{f} be the *self-consistent field* Fock operator and

$$\{|\mu\rangle\}_{\mu=1}^N \rightarrow \text{occupied orbitals}$$

$$\{|n\rangle\}_{n>N} \rightarrow \text{virtual orbitals}$$

Clearly $|n\rangle \in \{|\mu\rangle\}_{\mu=1, N}^{\perp}$. Let us put $|n\rangle = |\delta\mu\rangle$ in eq.(1). It follows

$$\langle \delta\Psi | H | \Psi \rangle = 0, \text{ with } |\delta\Psi\rangle = \sqrt{N!} A |1, 2, \dots n \dots N\rangle$$

⁷Note that each electron feels *all* the N electrons, *i.e.* this Hartree potential contains the so-called *self-interaction*. The latter cancels when the exchange term is taken into account.

⁸Note, in contrast, that for a local operator v we have $\langle \mathbf{x} | v | \mathbf{x}' \rangle = v(\mathbf{x}) \delta(\mathbf{x} - \mathbf{x}')$, where for $\mathbf{x} = (\mathbf{r}, \sigma)$ we have $\delta(\mathbf{x} - \mathbf{x}') = \delta(\mathbf{r} - \mathbf{r}') \delta_{\sigma\sigma'}$.

(with n in place of μ), i.e. *singly excited configurations do not mix with the HF one* (**Brillouin's theorem**)⁹.

Note 3¹⁰ A Slater determinant is *invariant* under non-singular (but otherwise arbitrary) linear transformation of vectors. That is, let

$$|\mu'\rangle = \sum_{\mu} T_{\mu'\mu} |\mu\rangle, \det T \neq 0$$

define a non-singular transformation and apply the multiplication rule of determinants. It follows

$$|\Psi'\rangle = |1'2'..N'\rangle = \det T |12..N\rangle$$

where the normalization factor $\det T$ (which is non null since T is non-singular) is immaterial for physical purposes. From a geometrical point of view, $\{|\mu\rangle\}_{\mu=1,N}$ span an N -dimensional linear space and T is a change of basis in this space. The (N -particle) HF vector depends *on the N -dimensional linear space **only***, i.e. it depends on the *projector* ρ in this space. For orthogonal orbitals it reads as

$$\rho = \sum_{\mu=1}^N |\mu\rangle \langle \mu|$$

and satisfies of course

$$\rho^2 = \rho = \rho^\dagger$$

This projector is the one-particle (HF) *density operator*. The above comment suggests that density operators of whatever order, *in the HF approximation*, depend on the 1-particle one **only**. This is indeed the case, in that it can be shown that for any order p

$$\rho^{(p)} = \underbrace{\rho \otimes \rho \otimes \rho \dots \otimes \rho}_{p \text{ times}} A_p$$

where

$$A_p = \frac{1}{p!} \sum_{P \in S_p} \epsilon_P P$$

⁹This is because we are imposing that the variation of the energy functional must be null at *first* order.

¹⁰This is a somewhat advanced topic, and can be skipped at a first reading.

For instance, for the second order (2-particle) density operator we have

$$\rho^{(2)} = \frac{1}{2!} \rho \otimes \rho (1 - \pi)$$

It follows that the HF equations can also be obtained by writing the expectation value of H in terms of the *one*-particle density matrix ρ and varying it under the *normalization constraint*

$$\text{Tr} \rho = N$$

and the projector properties

$$\rho^2 = \rho = \rho^\dagger$$

This results in the equation

$$[f[\rho], \rho] = 0$$

which is equivalent to the HF equations given above. Here, we emphasized the functional dependence of f on ρ which makes the problem self-consistent.